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#### 1 Evaluation of nitrous acid sources and sinks in urban outflow

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## 20 ABSTRACT

- Intensive air quality measurements made from June 22-25, 2011 in the outflow of the Dallas-Fort
- 22 Worth (DFW) metropolitan area are used to evaluate nitrous acid (HONO) sources and sinks. A
- two-layer box model was developed to assess the ability of established and recently identified
- 24 HONO sources and sinks to reproduce observations of HONO mixing ratios. A baseline model
- 25 scenario includes sources and sinks established in the literature and is compared to scenarios
- 26 including three recently identified sources: volatile organic compound-mediated conversion of
- nitric acid to HONO (S1), biotic emission from the ground (S2), and re-emission from a surface
- nitrite reservoir (S3). For all mechanisms, ranges of parametric values span lower- and upperlimit values. Model outcomes for 'likely' estimates of sources and sinks generally show under
- limit values. Model outcomes for 'likely' estimates of sources and sinks generally show under-prediction of HONO observations, implying the need to evaluate additional sources and
- variability in estimates of parameterizations, particularly during daylight hours. Monte Carlo
- simulation is applied to model scenarios constructed with sources S1-S3 added independently
- and in combination, generally showing improved model outcomes. Adding sources S2 and S3
- 34 (scenario S2/S3) appears to best replicate observed HONO, as determined by the model
- coefficient of determination and residual sum of squared errors ( $r^2 = 0.55 \pm 0.03$ , SSE =  $4.6 \times 10^6$
- $\pm 7.6 \times 10^5$  ppt<sup>2</sup>). In scenario S2/S3, source S2 is shown to account for 25% and 6.7% of the
- nighttime and daytime budget, respectively, while source S3 accounts for 19% and 11% of the
- nighttime and daytime budget, respectively. However, despite improved model fit, there remains
- 39 significant underestimation of daytime HONO; on average, a 0.15 ppt/s unknown daytime
- 40 HONO source, or 67% of the total daytime source, is needed to bring scenario S2/S3 into
- 41 agreement with observation. Estimates of 'best fit' parameterizations across lower to upper-limit
- values results in a moderate reduction of the unknown daytime source, from 0.15 to 0.10 ppt/s.
- 43 Keywords: air quality; unknown HONO source; Monte Carlo simulation; evolutionary solver

#### 44 1. INTRODUCTION

Atmospheric nitrous acid (HONO) is important due to the role of HONO in generation of 45 the hydroxyl radical (OH). There are a number of known sources of OH in the troposphere; 46 however, OH production from HONO is of interest because the sources, fate, and diurnal cycling 47 of HONO in the atmosphere have only recently begun to be elucidated. Models of atmospheric 48 HONO generally employ a mass balance approach that allows evaluation of the HONO budget, 49 50 often with a potentially limiting photostationary state assumption. As summarized by Spataro 51 and Ianniello (2014) models generally include sources, sinks, and transport, the last relevant as formation processes hypothesized to occur at the ground result in vertical gradients of HONO. 52 53 Homogeneous and heterogeneous reactions, as well as direct emission of HONO from combustion sources, contribute to the presence of HONO in the troposphere (Finlayson-Pitts and 54 Pitts, 1999). Nitrous acid strongly absorbs sunlight at wavelengths shorter than 390 nm resulting 55 56 in photolytic degradation to OH and nitric oxide (NO). This results in suppressed, but non-zero, mixing ratios of daytime HONO due to the presence of daytime sources (Kleffmann, 2007). At 57 night, the absence of this photolytic loss mechanism results in HONO accumulation, generally 58 on the order of 0.1 ppb to 10 ppb (Kleffmann et al., 2003; Su et al., 2008; Young et al., 2012). 59 The resumption of HONO photolysis after sunrise can lead to substantial formation of OH in the 60 early morning. Alicke et al. (2003) report that during the BERLIOZ investigation at a rural, 61 lightly trafficked site with low anthropogenic emissions during the summer months, photolysis 62 of HONO was the dominant source of OH in the morning, and contributed as much as 20% of 63 24-h integrated OH production. 64

Modeling studies generally show the need for an unknown daytime source to close the
HONO budget (Staffelbach et al., 1997; Lee et al., 2015). A number of photochemically driven

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homogeneous reactions have been identified or considered: e.g., the known reaction of OH and 67 NO and the hypothesized reaction of photolytically excited nitrogen dioxide (NO<sub>2</sub>) and water (Li 68 et al., 2008). The latter, however, may not proceed sufficiently rapidly or at adequate yields to 69 affect HONO mixing ratios in the atmosphere (Carr et al., 2009). Other potential homogeneous 70 sources are under discussion and review. For example, Li et al. (2014) proposed an internal 71 source of HONO that consumed nitrogen oxides, although follow up discussion and further 72 73 experiments indicate the source was likely strongly overestimated (Li et al., 2015; Ye et al., 74 2015).

Nitrous acid formation mediated by aerosol surface area (SA) is a topic of ongoing 75 76 research, largely because the complexity of aerosols results in substantial uncertainty regarding 77 their ultimate role in HONO formation. Static surfaces such as the ground (Stemmler et al., 78 2006) also may enhance HONO formation. Other hypothesized daytime sources include 79 emissions resulting from acid/base chemistry in soils (Su et al., 2011) and photolysis of nitric acid (HNO<sub>3</sub>) on forest canopy surfaces (Zhou et al., 2011). Photoenhanced conversion of NO<sub>2</sub> on 80 organic surfaces, including the ground and aerosols, are also thought to contribute to the daytime 81 HONO budget (George et al., 2005; Stemmler et al., 2006, 2007). 82

Given the many identified and proposed HONO source and sink mechanisms, single value estimates of parameterizations of HONO sources and sinks limit the ability to understand the impact of variability in multiple input parameters on models of HONO dynamics in the atmosphere. Monte Carlo simulation (MCS) provides a tool to observe the combined effects of ranges of input parameters and the resulting impact on the agreement between model output and measurements. In this work, we identify fourteen HONO sources or sinks established in the literature, including three sources that have recently (2013-2014) been identified. We evaluate

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these recently identified sources through incorporation into a baseline model with a full-factorial,
deterministic screening analysis. We then identify scenarios for which we stochastically
parameterize source and sink mechanisms with MCS to determine probability distributions of
modeled HONO mixing ratios.

#### 94 **2. METHODS**

95 2.1 Measurements

Measurements of gas- and particle-phase constituents were made from May 30 to July 1, 96 2011 in a semi-urban area approximately 68 km northwest of the Dallas-Fort Worth (DFW) 97 metropolitan area. The monitoring site was co-located with the Texas Commission on 98 99 Environmental Quality Eagle Mountain Lake (EML) continuous ambient monitoring station (CAMS 75). Further details regarding the geography, surrounding industrial and biogenic 100 activities, and site conditions have been outlined previously (Rutter et al., 2015) 101 102 Temperature, humidity (Vaisala, HMP-45C in a RM Young 10-plate solar radiation shield), and planetary boundary layer (PBL) height (Vaisala, CL31) were measured throughout 103 the duration of the campaign. Mixing ratios of HONO and HNO<sub>3</sub> were measured every five 104 minutes using a method that coupled a mist chamber with ion chromatography (Dionex, CD20-105 1), described in greater detail elsewhere (Dibb et al., 2004). First-order photolysis rate constants 106 (*j*-values) were determined with radiometric measurements of actinic flux determined with a 2-pi 107 108 double monochrometer with photomultiplier and subsequent calculations following IUPAC recommendations. Nitrogen oxides were recorded every minute using a chemiluminescence trace 109 level NO-NO<sub>2</sub>-NO<sub>x</sub> analyzer (Thermo Electron Corp., Model 42C) equipped with a Blue Light 110 Converter (Air Quality Design, Inc.) for NO<sub>2</sub> quantification. Hydroxyl radical was observed 111 using atmospheric pressure chemical ionization mass spectrometry (Kim et al., 2013). One-hour 112

113	averaged mixing ratios of volatile organic compounds (VOCs) were measured using a thermal
114	desorption gas chromatograph with flame ionization detection (Perkin-Elmer O <sub>3</sub> Precursor
115	Analyzer System). Continuous measurements of number-based particle size distributions
116	(diameter range of 20 nm to 500 nm) were made every ten minutes with a scanning electrical
117	mobility sizer (SEMS, Brechtel Inc. Model 2002) and were converted to SA distributions
118	assuming spherical particles. Concentrations of particulate phase nitrate were determined with an
119	Aerodyne high-resolution time-of-flight aerosol mass spectrometer, as described by Rutter et al.
120	(2015). Black carbon concentrations were measured using an aethalometer.

#### 121 2.2 Baseline model

122 A two-layer box model describing HONO mixing ratios was developed, with the height of the first layer set to 36 m to represent a surface layer and the height of layer 2 set to 72 m to 123 facilitate use of HONO observations above the surface layer that are available in the literature. 124 Established source (labeled as 'B1-B8' in Table 1) and sink mechanisms (labeled 'L1-L3' in 125 Table 1) are described in full in the Supporting Information (SI) (including Figures S1-S5 and 126 127 equations S1-S20). The timeframe selected for continuous modeling was 22 June 01:00 to 25 June 14:00 (all times local) based on the longest uninterrupted period during the campaign with 128 observations of HNO<sub>3</sub>, HONO, aerosol SA, NO<sub>2</sub>, NO, gas-phase chloride (assumed to be 129 hydrochloric acid, HCl), and  $j_{HONO}$ . Mixing ratios of constituents during this period were 130 generally typical of the broader study period. Equation 1 describes baseline sources and sinks 131 modeled with a transient approach: 132

$$\frac{d[\text{HONO}]_{trans}}{dt} = F_{B1} + F_{B2} + F_{B3} + F_{B5} + F_{B6} + F_{B7} + F_{B8} - (F_{L1} + F_{L2} + F_{L3}) - \Psi_{trans}$$
(1)

where [HONO]<sub>*trans*</sub> is the mixing ratio of HONO from modeled transient sources and sinks (ppt), *dt* is the time step (s) between measurements for which observations of all constituents present in Equation 1 were made, *F* represents the source or sink strength of the indicated mechanism (ppt/s), and  $\Psi_{trans}$  is the loss (or source) of HONO from layer 1 to (or from) layer 2 due to vertical transport (ppt/s).

Equation 1 describes the transient processes occurring in the model; source B4 wasincorporated into the model after accounting for transient processes as shown in Equation 2:

$$[\text{HONO}]_{\text{total}} = [\text{HONO}]_{\text{trans}} + f_{\text{emiss}}\Delta[\text{NO}_{x}]$$
(2)

where [HONO]<sub>total</sub> is the mixing ratio of HONO at a time step resulting from transient and
instantaneous processes (ppt) and *f<sub>emiss</sub>* is the direct HONO emission factor described in Table 1.
Equation 2 may overestimate the contribution of B4 in a box-model, as during the daytime,
HONO will rapidly photolyze prior to the measurement of emitted NO<sub>x</sub>.

144 Vertical transport,  $\Psi_{\text{trans}}$  (ppt/s), is calculated using a first-order flux-gradient relationship 145 simulated with the 1D CACHE model (Bryan et al., 2012) where mass is transported by eddy 146 diffusion at a magnitude proportional to the eddy diffusivity for heat ( $K_h$ ), shown in equation 3:

$$\Psi_{trans} = -K_h(z,t) \frac{\partial C(z,t)}{\partial z} \frac{1}{h}$$
(3)

147 where  $K_h(z,t)$  is the eddy diffusivity (m<sup>2</sup>/s) at height z (m) and time t. As shown in equation 3, 148 estimates of flux are divided by h, the height of the second layer in the model (m), prior to 149 inclusion in equation 1.

Two 1D simulations during the campaign were used to derive  $K_h$ , including one simulation for 7-9 June and one for 10-12 June. For the layers corresponding to the upper boundary that are used in the results here,  $K_h$  is derived based on a length scale, vertical wind

153	shear, and a stability parameter (Forkel et al., 1990). It is calculated at each time step within the
154	model, providing a diurnal cycle that is based on meteorological conditions during the campaign.
155	Observations of HONO were made at one elevation, approximately 10 m above surface,
156	and were used to represent the HONO mixing ratio in layer 1 of the model. Equation 3 requires
157	an estimate of the HONO mixing ratio in layer 2 to estimate the HONO gradient. Three scenarios
158	were considered: 1) no gradient (i.e., [HONO] in layer 1 equals that in layer 2 at all times); 2) a
159	gradient created using fractions of [HONO] presented in Vandenboer et al. (2013), representative
160	of a stronger nighttime gradient and a weaker daytime gradient (GrN); and 3) a gradient created
161	from fractions of [HONO] presented in Villena et al. (2011) that is representative of a stronger
162	daytime gradient and weaker nighttime gradient (GrD). Diurnal profiles of the three gradient
163	conditions are shown in Figure S6 of the SI and implications of this limitation are discussed in
164	Section 3.2.
165	2.3 Parameterization and evaluation of newly identified HONO sources
166	Three recently identified HONO source mechanisms were parameterized to assess the
167	potential of these mechanisms (in conjunction with B1-B8 and L1-L3) to independently or
168	jointly account for HONO mixing ratios observed in DFW. The three mechanisms, listed in
169	Table 1 as S1, S2, S3 are incorporated into Equation 1 as additional sources of HONO.
170	Source S1 is the formation of HONO from the reduction of $HNO_3$ to HONO mediated by
171	VOCs emitted from motor vehicles (Rutter et al., 2014). The source strength ( $F_{S1}$ , ppt/s) was
172	parameterized using HONO source strength and reactant mixing ratios presented in Table 1 of
173	Rutter et al. (2014) and is shown in equation 4:

$$F_{S1} = f_{HNO_3,VOC} \left[ \frac{\left( \frac{[\text{Propylene}]}{[\text{Benzene}]} \right)_{EML}}{\left( \frac{[\text{Propylene}]}{[\text{Benzene}]} \right)_{Max,DFW}} \right] \left[ \frac{[\text{HNO}_3]_{EML}}{[\text{HNO}_3]_{rutter}} \right]$$
(4)

where  $f_{HNO3,VOC}$  is the observed HONO formation rate (ppt s<sup>-1</sup>) in Rutter et al. (2014), and 174 normalizing ratios are further described in the SI. Estimates of 'likely' f<sub>HNO3,VOC</sub> were taken for 175 experiments conducted at 50% RH while 'lower-limit' and 'upper-limit' estimates were taken as 176 the minimum and average across experiments shown in Table 1 of Rutter et al. (2014). 177 Normalizing assumptions shown in equation 4 resulted in, on average, ~95% reduction of 178  $f_{HNO3,VOC}$  when calculating  $F_{S1}$ . The form of the parameterization in equation 4 is speculative; 179 180 propylene is chosen as a proxy for reactive VOCs while benzene is chosen to account for dilution that may occur as air masses move from DFW to EML (see Figure S7 in the SI for a diurnal 181 profile of propylene/benzene). Identification of specific reactive species participating in the 182 HONO formation process identified in Rutter et al. (2014) would enable improvements in 183 development and assessment of parameterizations of VOC-mediated conversion of HNO<sub>3</sub> to 184 185 HONO.

186 Source S2 is HONO emissions from soil bacteria as described by Oswald et al. (2013). 187 Emission from the soil ( $F_{S2}$ , ppt/s) was assumed to mix instantaneously through the first model 188 layer as shown in equation 5:

$$F_{S2} = \frac{f_{soil}}{h} \Gamma_{S2}$$
(5)

189 where  $f_{soil}$  is the "optimum" HONO flux from a soil type (molec cm<sup>-2</sup> s<sup>-1</sup>), *h* is the height of the 190 model layer, and  $\Gamma_{S2}$  represents the conversion factor to ppt/s prior to inclusion in equation 1 (see 191 the SI equations S21-S24 for an example calculation). The 'lower-limit' value of  $f_{soil}$  was taken

192	as the value of HONO flux for pasture, and the 'upper-limit' value was taken as that for
193	grassland. No 'likely' value of $f_{soil}$ was selected, as pasture and grassland were the only two
194	relevant soil types for the DFW region. Despite specifying a 'lower-limit' value, this
195	investigation may be effectively considering the high end of contribution of soil bacteria to
196	HONO because "optimum" values of flux are used for both soil types.

Source S3 is the re-emission of HONO from a surface nitrite reservoir by displacement
from HNO<sub>3</sub> and HCl, as in Vandenboer et al. (2014, 2015) and shown in equation 6:

$$F_{S3} = \frac{[\text{HNO}_3] + [\text{HCl}]}{h} v_d \eta$$
(6)

where  $F_{S3}$  is the source strength of S3 (ppt s<sup>-1</sup>),  $v_d$  is the deposition velocity of HNO<sub>3</sub> and HCl, 199 taken as 1 cm s<sup>-1</sup>, and  $\eta$  is the displacement efficiency, ranging from 1% to 9% to 20% for 200 201 'lower-limit', 'likely', and 'upper-limit' values, respectively (VandenBoer et al., 2014). This parameterization was constrained by the calculation of a 'reservoir' of nitrite from deposited 202 HONO, approximated from a material balance on the ground where the source of nitrite is 203 mechanism L1 and loss is due to displacement from mechanism S3. Mechanism S3 was set to 0 204 when the reservoir was equal to 0. As there may be additional sources of surface nitrite other 205 206 than gas-phase HONO and surface nitrite accumulation over greater than diurnal time-scales, 207 equation 6 likely represents a conservative estimate of the source strength of S3. Further description of the constraints on source S3 is given in the SI and dynamics are depicted in Figure 208 S8, also in the SI. 209

210 2.4 Model calculation and assessment

Nitrous acid mixing ratios were first modeled with the baseline scenario using the B and
L parameterizations summarized in Table 1. The 'likely' parameterization incorporates HONO

source and sink estimations thought most representative of each mechanism, while 'upper-limit' 213 and 'lower-limit' are values that result in maximum or minimum HONO production, 214 respectively, e.g. in the 'upper-limit', parameterizations of sources result in greater formation 215 while those of sinks result in lower loss rates. Predictions of HONO mixing ratios were assessed 216 through the residual sum of squared errors (SSE) and the coefficient of determination  $(r^2)$ , both 217 determined from differences between modeled and measured HONO mixing ratios. 218 219 Model scenarios were constructed to assess the three new mechanisms (mechanism ID = 220 S1, S2, and S3 shown in Table 1) and gradient conditions (GrN or GrD); scenarios are named according to the gradient used and sources added, e.g., GrN S2/S3 refers to a model scenario 221 with the stronger nighttime gradient as described previously and with sources S2 and S3 added to 222 baseline sources B1-B8 and sinks L1-L3. Sources S1-S3 were added to the baseline model in a 223 full-factorial deterministic screening analysis (using 'likely' estimates of parameterizations) to 224 225 identify scenarios for further analysis. Monte Carlo simulation (Crystal Ball v. 11.1.2.3, Oracle) was used to evaluate the probability of model scenarios to account for observed HONO mixing 226 ratios. Input distributions of source and sink parameterizations were assumed to be triangular 227 probability distributions, bounded by 'lower-limit' and 'upper-limit' values with the 'likely' 228 value as the most frequently occurring. Model sensitivity to the number of trial simulations was 229 performed to ensure a trial-independent solution was achieved; all MCS were conducted with 230 5,000 iterations. A bounded evolutionary solver was applied to the baseline model scenario and 231 to the model scenario with the highest  $r^2$  and lowest residual SSE in the deterministic screening 232 analysis. The evolutionary solver used a genetic algorithm to estimate source and sink 233 parameterizations with a minimum SSE across the range of 'lower-limit' to 'upper-limit' values 234 for each source or sink mechanism. 235

#### 236 **3. RESULTS AND DISCUSSION**

#### 237 *3.1 Ambient air monitoring in the outflow of DFW*

Experimental observations of mixing ratios of ambient gases and particles input to the 238 model are shown in Figure 1; diurnal profiles of selected constituents across the full monitoring 239 campaign are shown in Figure S9 of the SI. Values of HONO/NO<sub>2</sub> are variable and elevated 240 during the daytime, possibly indicative of a secondary daytime source of HONO. Mixing ratios 241 242 of HNO<sub>3</sub> are suppressed in the morning and evenings and elevated during daytime hours, likely a 243 result of strong daytime HNO<sub>3</sub> production from the reaction of NO<sub>2</sub> and OH (Aneja et al., 1994). The highest observed mixing ratios of HNO<sub>3</sub> across the full monitoring campaign are included in 244 245 the model period shown in Figure 1, exceeding 5000 ppt in the early evening of June 22, 2011. Mixing ratios of HCl exhibit similar trends to those observed for HNO<sub>3</sub>. Mixing ratios of HONO 246 show accumulation over the nighttime and suppression during the daytime, a result of the strong 247 248 loss due to photolysis and convective dilution during the daytime hours. Aerosols and aerosolphase constituents appear elevated during the nighttime hours of 6/23 and 6/24 compared to 249 daytime concentrations, but are suppressed during the nighttime of 6/25. Across the model 250 period, the SA of particulate matter averages  $125 \,\mu\text{m}^2 \,\text{cm}^{-3}$ , consistent with typical values across 251 the month-long monitoring campaign (Figure S1), and ranges  $22 \,\mu m^2 \, cm^{-3} - 392 \,\mu m^2 \, cm^{-3}$ . 252

*3.2 Baseline model* 

Mixing ratios of HONO are first calculated with the model under the baseline scenario for 'likely' estimates of parameterizations. Predicted and measured mixing ratios of HONO for the baseline scenario with three HONO gradient conditions described in Section 2.2 are shown in Figure 2. The "no gradient" condition results in substantial over-estimation of nighttime HONO mixing ratios, logical given the role of the ground surface in HONO formation processes

259	included in the baseline scenario and the first layer height of 36 m. Conversely, the GrN and GrD
260	conditions both result in underestimation of nighttime HONO, with relatively small differences
261	between the two conditions. A strong daytime sink, due to photolysis, results in suppression of
262	modeled daytime mixing ratios below observation for all three gradient conditions, implying the
263	need for daytime sources beyond those considered in the baseline scenario. The underestimation
264	may also result from the limited vertical resolution in the two-layer box model used here and the
265	measurement height in the lower portion of the first layer (10 m); it is likely that a continuous
266	HONO gradient is present in the 36 m of the model first layer resulting in a lower modeled
267	mixing ratio across the first model layer than the 10 m observation.
268	While relatively few studies report measurements of vertical gradients of HONO
208	while relatively lew studies report measurements of vertical gradients of fronto,
269	available profiles generally show higher HONO mixing ratios in surface layers than aloft,
270	indicative of ground surface HONO formation. Michoud et al. (2014) summarize several studies
271	reporting vertical gradients, four of which show the presence of a vertical gradient (Veitel, 2002;
272	Zhang et al., 2009; Villena et al., 2011; Wong et al., 2012) and one that does not (Häseler et al.,
273	2009). Vandenboer et al. (2013) report high-resolution vertical profiles measured from a tower in
274	Boulder, CO, and show the presence of both daytime and nighttime HONO gradients. Veitel et
275	al. (2002) report that over 13 months of measurements, HONO mixing ratios were observed to
276	decrease with height under nearly all atmospheric conditions. For the present investigation, we
277	interpret the over-prediction of HONO mixing ratios in the nighttime for the "no gradient"
278	condition, when convective mixing is most likely to be diminished, to indicate a HONO vertical
279	gradient. Thus, conditions GrN or GrD better represent the vertical structure of HONO mixing
280	ratios in the outflow of DFW. While this appears to be in agreement with the preponderance of
281	available HONO vertical gradient measurements, a site-specific HONO gradient would clearly

improve the present study. Nevertheless, parameterizations here allow an estimation of the
source and sink processes in the outflow of DFW and exploration of two estimates of gradients
to assess model sensitivity to the HONO vertical profile. The impact of the vertical gradient and
of parameterizations of established and recently identified HONO sources and sinks are further
explored in Sections 3.3-3.5.

#### 287 *3.3 Deterministic screening analysis*

A deterministic screening analysis was employed to evaluate model outcomes when sources S1-S3, acting independently or in any combination, are incorporated into the model. This full-factorial analysis, consisting of 24 possible scenarios, is conducted for only the 'likely' parameterizations of the mechanisms, as shown in Table S1 of the SI. Full output of model runs across all gradient conditions and scenarios of parameterizations are provided in Figures S10-S12.

Generally, 'likely' estimates of parameterizations showed improved model fit compared 294 to 'upper-limit' estimates, implying additional sources of HONO, rather than increased 295 296 production from baseline sources result in improved model outcomes. Subsequent discussion in this section reflects 'likely' parameterizations. Scenarios identified for further investigation are 297 those with a combination of low SSE and high  $r^2$ . The baseline model generally is characterized 298 by the highest model SSE, and the addition of source mechanisms S1-S3 generally lowers SSE 299 and increases  $r^2$ . In cases, however, the SSE is lowered while the  $r^2$  decreases (for example, from 300 GrN Baseline to GrN S1). This is a result of improvement in model prediction for only a subset 301 302 of times in the modeling period. The screening analysis identified scenario S2/S3 and scenario S1/S2/S3 as having the lowest SSE and highest  $r^2$  (SSE range:  $4.3 \times 10^6 - 6.7 \times 10^6$ ;  $r^2$  range: 0.42-303

304 0.58). These scenarios, along with baseline scenarios for comparison, are further explored with305 MCS and an evolutionary solver.

#### 306 *3.4 Monte Carlo simulation*

Six model scenarios that vary the new sources and vertical gradient conditions were 307 evaluated with MCS to incorporate uncertainty and variability in each mechanism into the 308 model; model estimates of HONO are determined as probabilistic distributions at each model 309 time step. Summarized output of MCS are shown in Figure 3 as hourly-averaged diurnal profiles 310 of measured and modeled distributions of HONO mixing ratios across the model period. The 311 MCS reinforces the conclusions that 'baseline' source mechanisms cannot explain observed 312 HONO mixing ratios; in the GrN Baseline condition, 90<sup>th</sup> percentile values of model output 313 underestimate observed HONO mixing ratios in 23 of 24 reported hours, and 75<sup>th</sup> percentile 314 values underestimate observed HONO mixing ratios all 24 reported hours. 315

The addition of source mechanisms S2 and S3 to the model (Figure 3) results in 316 improved agreement between the model and observations for nighttime mixing ratios of HONO 317 for both GrN and GrD conditions. GrN S2/S3 shows 9 of the 10 hours in the 21:00-07:00 318 nighttime period are between the 10<sup>th</sup> and 90<sup>th</sup> percentile values determined in the model. GrD 319 S2/S3 shows improvement over the GrD Baseline condition; however, metrics of goodness of fit 320 are lower than GrN S2/S3, and there is less improvement over baseline. This appears to be a 321 result of sustained accumulation over the nighttime period, due to the smaller HONO nighttime 322 vertical gradient in the GrD condition. Under both GrN and GrD conditions for scenario S2/S3, 323 324 daytime mixing ratios of HONO remain substantially underpredicted as in the baseline condition.

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325	The addition of all three sources (S1, S2, and S3) does not appear to resolve
326	underprediction of the daytime HONO mixing ratio. In the GrN condition, the addition of source
327	S1 results in a small increase in over-estimation of nighttime HONO mixing ratios and metrics of
328	model fit worsen. In the GrD condition, there is a limited impact from the combined effect of
329	sources S1, S2 and S3, with a modest reduction in both SSE and correlation coefficient when
330	comparing GrD S1/S2 to GrD S1/S2/S3. Figure 3 shows GrN S2/S3 results in improved model
331	fit compared to other scenarios, although daytime HONO remains substantially underestimated.
332	An estimation of average total and relative source and sink strength across both nighttime
333	(21:00 - 07:00) and daytime $(07:00 - 21:00)$ is shown in Figure 4 for GrN S2/S3. Estimates of
334	sources and sinks are reported for 'likely' values of parameterizations for the indicated time
335	period. Considerable temporal differences in the contributions of various source and sinks to the
336	HONO budget exist. At night, HONO from NO <sub>2</sub> conversion at the ground (B7) is the major
337	source, contributing 53% of the HONO budget. Biotic release from the ground (S2) and re-
338	emission from the nitrite reservoir (S3) are the next two largest contributors at 25% and 19%,
339	respectively. Nighttime HONO is slightly over-estimated; an 'unknown' nighttime sink of
340	0.0016 ppt/s, or 3% of the total, is required to bring the model into agreement with observations.
341	Major nighttime sinks are vertical transport and deposition of HONO at the ground surface,
342	contributing 73% and 21%, respectively. These nighttime sources and sinks are in general
343	agreement with relative estimates of mechanisms reported by Czader et al. (2012), who report
344	71% of HONO production due to heterogeneous surface chemistry and losses due to transport
345	and deposition of 77% and 23%, respectively, during the nighttime and pre-sunrise morning.
346	During the daytime, a missing HONO source dominates; however there are meaningful
347	contributions to the daytime HONO budget from S3, S2, B8, B7 and B5. A missing daytime

source of 0.15 ppt s<sup>-1</sup>, or 67% of the total HONO source budget shown in Figure 4, is needed to 348 bring modeled and measured results into full agreement. This "missing" source is in the range of 349 magnitudes identified in other investigations, ranging from 0.03 - 0.3 ppt s<sup>-1</sup> (Su et al., 2008; 350 Elshorbany et al., 2009; Sörgel et al., 2011; VandenBoer et al., 2013; Lee et al., 2015). Unless 351 there is a positive artifact that depends on sunlight, a strong daytime source is needed to balance 352 the substantial sink of HONO due to photolysis (89% of the total sink). In section 3.5, we 353 explore the potential for 'best fit' estimates of parameterizations in GrN S2/S3 to close some 354 portion of the HONO budget through optimization of parameterizations across the range of 355 values presented in Table 1. 356

#### 357 *3.5 Evolutionary solver and sensitivity analysis*

An evolutionary solver was employed to estimate the optimal combination of input values within 'lower-limit' to 'upper-limit' ranges of parameterizations and the resulting impact on the estimate of the "missing" HONO source or sink. The evolutionary solver was applied to the GrN baseline scenario and GrN S2/S3. Model outcomes with optimal estimates for GrN baseline and GrN S2/S3 are shown in Figure 5 and parameterizations are reported in Table 2.

Across optimization of both GrN Baseline and GrN S2/S3, the largest changes to the 363 parameterizations relate to heterogeneous conversion of NO<sub>2</sub> on aerosol (B1 and B2) and on the 364 ground (B7, B8), and HONO uptake to the ground (L1). Aerosol processes increase substantially 365 as a result of a speculative upper-limit as described in the SI; B1 was allowed to vary over 1.5 366 orders of magnitude and B2 over 2.5 orders of magnitude based on prior modeling studies, rather 367 368 than experimental estimates. However, contributions from B1 and B2 remain limited (<1% as can be determined from absence of B1 and B2 in Figure 4), in part a result of the two layer box-369 model used here that emphasizes ground-level phenomena. In both GrN Baseline and GrN 370

S2/S3, the optimization resulted in B8 at the upper-limit of the parameterization. Source B7
increased by ~2× in GrN Baseline, but more moderately in GrN S2/S3, a result of the
contribution of sources S2 and S3 in GrN S2/S3. In GrN S2/S3, deposition loss (L1) increased, a
result of the need to balance increases in parameterizations of sources that act over both daytime
and nighttime periods (e.g., S3) and contribute to reductions in the daytime "unknown" source
but also nighttime accumulation.

377 Figure 5 shows greater improvements in metrics of model goodness of fit for the optimal solution of GrN S2/S3 compared to the optimal solutions of the GrN Baseline. This indicates that 378 baseline mechanisms are not able to similarly explain HONO observations under any 379 combination of input parameters compared to the scenario with S2/S3 present. This appears to 380 largely result from stronger parameterizations of S2/S3 resulting in improved estimates of 381 daytime HONO mixing ratio, although levels are still lower than observed. Best-fit 382 383 parameterizations of GrN S2/S3 result in a missing daytime source of 0.10 ppt/s, reduced from 0.15 ppt/s (Figure 4), implying that a substantial missing HONO source remains even across a 384 statistically optimized range of parameterizations. 385

The "best-fit" estimates of GrN S2/S3 reflect an improved statistical outcome for the 386 model when parameterizations are allowed to vary across a range of values. Parameterizations in 387 Table 2 with larger percentage changes imply a combination of model sensitivity to the 388 parameter as well as uncertainty in the value of the parameterization. We conducted a sensitivity 389 analysis to identify the most important parametrizations impacting the estimates of goodness-of-390 fit, the model  $r^2$  and SSE. The sensitivity analysis for GrN S2/S3 is summarized in Table S2 of 391 the SI, reported as the Spearman's rank correlation coefficient (p) between each mechanism's 392 input parameter and the model output  $r^2$  or SSE. Uptake of NO<sub>2</sub> at the ground (B7) is the 393

parameter with the largest impact on both the model SSE and  $r^2$ , by a comparatively large 394 margin. Given that there is a wide range of estimates of the uptake coefficient parameterizing B7 395 in the literature, this source represents a large source of uncertainty in the model. Sources S3, B8, 396 and S2 are the next three strongest correlations with model SSE; interestingly, all four sources 397 with highest sensitivity (B7, B8, S2, and S3) are ground-level phenomena. Source B7 was 398 strongest correlated with night-time (21:00-07:00) HONO mixing ratios while source S3 was 399 strongest correlated with daytime HONO. This underscores the importance of characterizing the 400 role of the ground surface mechanisms, including biotic release and ground-level chemical 401 transformations. 402

The presence of a substantial missing daytime source is further explored via estimation of 403 correlation coefficients between measured constituents and products of constituents with the 404 missing HONO source, similar to the analysis presented by Lee et al. (2015). This analysis 405 406 employed time-series measurements for constituents and the estimate of missing HONO at each time step required for model agreement with observation. Outcomes are shown in Table S3 for 407 'likely' and 'best-fit' estimates of GrN S2/S3. Relatively strong correlation coefficients ( $r^2 > 0.5$ ) 408 were observed for  $j_{NO2}$  and  $j_{NO2}$  × temperature with the missing HONO source, the latter in close 409 agreement to the results of Lee et al (2015). However, the correlation of  $j_{NO2} \times NO_2$  with the 410 missing HONO source is weak ( $r^2 = 0.09 - 0.17$ ), as is the correlation of  $j_{NO2} \times SEMS SA \times NO_2$ 411  $(r^2 = 0.08 - 0.16)$  and with NO<sub>2</sub> alone  $(r^2 = 0.21 - 0.25)$ . The stronger correlation with j<sub>NO2</sub> and 412  $j_{NO2}$  × temperature may imply photosensitized conversion on organics, including humic acids, 413 which are mainly ground surface sources (Stemmler et al., 2006, 2007), are underestimated. The 414 weak correlation of the missing HONO source with NO2 and products containing NO2 mixing 415 ratios appears aligned with a recent analysis of weekday-weekend HONO and NO<sub>2</sub> relationships 416

that shows HONO production rates do not increase with increases in NO<sub>2</sub>, implying daytime HONO production may not be rate-limited by NO<sub>2</sub> (Pusede et al., 2015). Weakening correlations for products of gas- and particle-phase constituents and  $j_{NO2}$  also may result from the two-layer model that lends greater emphasis to interactions at the ground level, consistent with the results of the sensitivity analysis in Table S2 and discussed previously.

422 3.6 Model limitations

423 The model described in this work is subject to a number of important limitations. Source S1 assumes the source strength determined in the laboratory is possible in the ambient 424 environment, with several normalizing assumptions. However, as we did not observe meaningful 425 426 formation of HONO from source S1, the impact of the speculative parameterization is therefore limited in this investigation. Future field efforts should further investigate the potential for VOC-427 mediated reduction of HNO3 to HONO in near-source environments. Source S2 was 428 429 parameterized using a single value for a model simulation; there are likely to be diurnal variations in biological activity and soil water content that would impact the parameterization of 430 source S2. Source S3 considered only gas-phase HONO as an input to the surface nitrite 431 reservoir and that the reservoir was empty at the beginning of the model period. This may result 432 in a conservative estimate of the contribution of source S3. 433

Input distributions in MCS were assumed to be triangular. This assumption may overweight estimates of parameterizations at the 'upper-limit' and 'lower-limit' extents of the distribution as compared to a normal distribution. A triangular distribution was chosen, in part, to ensure parameterizations did not exceed upper or lower-limit estimates in MCS. The two-layer box model uses instantaneous and *in-situ* mixing ratios to constrain the model, with the assumption of instantaneous mixing up to the first layer height. Transport between layers was

estimated using an approximation of HONO vertical gradients at similar heights taken from 440 literature. We assume transport time for  $NO_x$  sources that exceeds the atmospheric age of HONO 441 (Lee et al. 2013). During the daytime periods (07:00-21:00), the atmospheric age of HONO 442 across the modeling period in this work averaged 19.4 min and ranged from 8.9 to 128 min. We 443 assume NO<sub>x</sub> sources input to the model originate from the metropolitan DFW area (~70 km 444 away), while the wind speed averaged 19 km/h, resulting in a transport time of 220 min. 445

#### **4. CONCLUSIONS** 446

Model predictions of HONO that account for ranges in parameterizations of HONO 447 source and sink mechanisms enable a statistical assessment of the likelihood of the model to 448 449 match observation. Observations of HONO appear most accurately simulated when emission from soil biota (S2) and re-emission from a ground level nitrite source (S3) are included in the 450 model. Model output for GrN S2/S3 accounted for, on average, 33% of the daytime HONO 451 452 budget and 103% of the nighttime HONO budget. Major nighttime sources included (in order) NO<sub>2</sub> conversion at the ground (B7), biotic release from soil (S2), and re-emission from the nitrite 453 reservoir (S3). Major daytime sources include S3, S2, photoenhanced NO<sub>2</sub> conversion at the 454 ground (B8), B7, and the reaction of OH with NO (B5). Model fit improved after application of 455 an evolutionary solver, resulting in a reduction of the estimate of the unknown daytime source 456 for GrN S2/S3. However, the presence of a substantial unknown daytime source (on average 457 0.10 ppt/s) even with a statistically optimal fit for GrN S2/S3 implies additional sources of 458 HONO than those evaluated here must be included to reproduce accurately daytime HONO 459 mixing ratios. Analyses of model sensitivity and correlations between the missing HONO source 460 and constituents imply the presence of additional, or underestimation of considered, ground-level 461 HONO sources in this investigation. 462

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Mechanism	Ш	Parameter	Lower-limit	Likelv	Upper-limit	Reference
Aerosol uptake of NO <sub>2</sub>	B1	γ <sub>NO2</sub> (-)	$2.0 \times 10^{-7}$	$1.0 \times 10^{-6}$	$5.0 \times 10^{-6}$	Kleffmann et al. (1998); Aumont et al.(2003)
Photoenhanced aerosol uptake of NO <sub>2</sub>	B2	γ <sub>NO2,hv</sub> (-)	$4.0 \times 10^{-6}$	$1.0 \times 10^{-5}$	$1.0 \times 10^{-3}$	Stemmler et al. (2007); Wong et al. (2013)
Photoenhanced conversion of NO <sub>2</sub> soot	B3	$\frac{\gamma_{\text{soot,BET}}(-)}{\text{BET surface area (cm2/g)}}$	$4.0 \times 10^{-7}$ 9.7×10 <sup>5</sup>	$5.0 \times 10^{-7}$ $1.2 \times 10^{6}$	$\frac{6.0 \times 10^{-7}}{1.3 \times 10^{6}}$	Monge et al. (2010)
Direct HONO emission	B4	$f_{\rm emiss}$ (%v, $\Delta$ HONO/ $\Delta$ NO <sub>x</sub> )	0.0029	0.0055	0.0080	Kirchstetter et al. (1996); Kurtenbach et al. (2001)
	D5	$k \infty$ (T) (cm <sup>3</sup> molec <sup>-1</sup> s <sup>-1</sup> )	$3.0 \times 10^{-11}$	$3.6 \times 10^{-11}$	$4.3 \times 10^{-11}$	NASA (2011)
OH + NO	БЭ	$k_o(T) \ (cm^6 \ molec^{-2} \ s^{-1})$	$5.8 \times 10^{-31}$	$7.0 \times 10^{-31}$	$8.4 \times 10^{-31}$	NASA (2011)
HONO from surface	DC	$j_{\rm HNO3-HONO} (s^{-1})$	$1.0 \times 10^{-5}$	$1.2 \times 10^{-5}$	$1.4 \times 10^{-5}$	Zhou et al. (2003)
HNO <sub>3</sub> photolysis	во	$v_{d, HNO3} (cm s^{-1})$	1.50	1.75	2.25	Walcek et al. (1986)
HONO from NO <sub>2</sub> conversion at ground	B7	γ <sub>NO2, gr</sub> (-)	$1.0 \times 10^{-6}$	$5.0 \times 10^{-6}$	1.0× 10 <sup>-5</sup>	Kleffmann et al. (1998); Kurtenbach et al. (2001)
Photoenhanced NO <sub>2</sub> conversion, ground	B8	$\gamma_{\rm NO2,gr,hv}$ (-)	$1.7 \times 10^{-5}$	$2.0 \times 10^{-5}$	$6.0 \times 10^{-5}$	Stemmler et al. (2006); Wong et al. (2013)
$HNO_3 \rightarrow HONO, VOC$	<b>S</b> 1	$f_{HNO3, VOC}$ (ppt s <sup>-1</sup> )	$3.6 \times 10^{-2}$	$5.8 \times 10^{-2}$	$8.3 \times 10^{-2}$	Rutter et al. (2014)
Biotic release, ground	S2	$f_{soil}$ (molec cm <sup>-2</sup> s <sup>-1</sup> )	-	$1.7 \times 10^{9}$	$4.0 \times 10^{9}$	Oswald et al. (2013)
Re-emission from NO <sub>2</sub> -(p) reservoir	<b>S</b> 3	<i>v<sub>d</sub></i> ×η (cm s <sup>-1</sup> )	$1.0 \times 10^{-2}$	$9.0 \times 10^{-2}$	$2.0 \times 10^{-1}$	Vandenboer et al. (2014)
HONO uptake at ground	L1	γ <sub>HONO,gr</sub> (-)	$1.0 \times 10^{-4}$	$2.0 \times 10^{-5}$	$1.8 \times 10^{-5}$	Vandenboer et al. (2013); Wong et al. (2013); Trick (2004)
HONO + OH	L2	$k_{\text{HONO+OH}} (\text{cm}^3 \text{molec}^{-1} \text{s}^{-1})$	$6.75 \times 10^{-12}$	$4.5 \times 10^{-12}$	$3.0 \times 10^{-12}$	NASA (2011)
HONO photolysis	L3	$j_{\rm HONO}~({\rm s}^{-1})$	1.8	$\times 10^{-3} - 3.9 \times 10^{-3}$	10 <sup>-5 a</sup>	This investigation

#### 595 Table 1. HONO source and sink mechanisms considered for modeling HONO in the outflow of the DFW metropolitan area.

<sup>a</sup>Maximum-minimum range of the experimentally determined time-series values of  $j_{HONO}$  input to the model (not varied).

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597	Table 2. Best estimates of parameterizations of sources and sinks of HONO in the outflow of
598	DFW for baseline and scenario GrN S2/S3.

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		Best-fit estimate (% difference from 'likely')		
ID	Parameter	GrN S2, S3	GrN Baseline	
B1	γ <sub>NO2</sub> (-)	3.9 × 10 <sup>-6</sup> (294%)	$2.5  imes 10^{-6} (152\%)$	
B2	γ <sub>NO2,hv</sub> (-)	8.5 × 10 <sup>-4</sup> (8500%)	1.0 × 10 <sup>-3</sup> (9900%)	
D2	$\gamma_{soot,BET}$ (-)	5.3 × 10 <sup>-7</sup> (6%)	$5.3  imes 10^{-7} (7.1\%)$	
Ъ5	BET surface area (cm <sup>2</sup> /g)	$1.1 \times 10^2$ (-6.5%)	$1.2 \times 10^2$ (-3%)	
B4	$f_{\rm emiss}$ (%v, $\Delta$ HONO/ $\Delta$ NO <sub>2</sub> )	0.0043 (-22%)	0.0049 (-10%)	
D5	$k\infty(T)$ (cm <sup>3</sup> molec <sup>-1</sup> s <sup>-1</sup> )	$3.7 \times 10^{-11} (4.4\%)$	$3.8  imes 10^{-11}$ (4.8%)	
ЪЭ	$k_o(T) \ (cm^6 \ molec^{-2} \ s^{-1})$	$7.6  imes 10^{-31} (9\%)$	$7.3  imes 10^{-31}$ (4.8%)	
D6	$j_{ m HNO3-HONO}~( m s^{-1})$	$1.2 \times 10^{-5} (-3\%)$	$1.3 \times 10^{-5} (7.7\%)$	
р0	$v_{d, HNO3} (cm s^{-1})$	1.8 (4.6%)	2.0 (17%)	
B7	γ <sub>NO2, gr</sub> (-)	6.1 × 10 <sup>-6</sup> (22%)	9.9 × 10 <sup>-6</sup> (97%)	
B8	$\gamma_{NO2,gr,hv}$ (-)	$6 \times 10^{-5} (200\%)$	$6  imes 10^{-5} (200\%)$	
<b>S</b> 1	$f_{HNO3, VOC}(\text{ppt s}^{-1})$	n/a	n/a	
S2	$f_{soil} (\mathrm{molec} \mathrm{cm}^{-2} \mathrm{s}^{-1})$	$2.8 \times 10^{9}$ (66%)	n/a	
<b>S</b> 3	$v_d \times \eta$ (cm s <sup>-1</sup> )	0.18 (105%)	n/a	
L1	γ <sub>HONO,gr</sub> (-)	5.7 × 10 <sup>-5</sup> (185%)	2.0×10 <sup>-5</sup> (-1.1%)	
L2	$k_{\text{HONO+OH}} (\text{cm}^3 \text{molec}^{-1} \text{s}^{-1})$	$5.7 \times 10^{-12} (28\%)$	$4.6 \times 10^{-12} (2.1\%)$	
L3	$\dot{j}_{\mathrm{HONO}}~(\mathrm{s}^{-1})$	unchanged	unchanged	
	Missing source or sink: daytime, nighttime (ppt s <sup>-1</sup> )	0.10, -0.0112	0.15, -0.006	

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Figure 1. Time series inputs to the two-layer box model of HONO mixing ratios in the outflow of DFW.



Figure 2. Model output for 'likely' estimates of parameterizations under conditions of no gradient, stronger nighttime gradient (GrN), and stronger daytime gradient (GrD).



Figure 3. Summary of Monte Carlo simulation output for baseline scenarios, and scenarios with S2/S3 and S1/S2/S3 added to the baseline scenario.



Figure 4. Relative contribution to HONO source or sink strength in GrN S2/S3 with 'likely' estimates of parameterizations. Contributions are averaged for the time period indicated above each pie chart across the modeling period (6/22/2011 01:00 – 6/25/2011 14:00 local time). Unknown source or sink is determined by stepwise addition of HONO source or sink such that modeled HONO equals measured HONO.

■ B8, Photoenhanced conversion of NO<sub>2</sub> at ground



Figure 5. Model performance with best-fit parameters for the nighttime gradient (GrN) scenario with sources S2 and S3, compared to the nighttime gradient scenario with only baseline sources included.

- A two-layer box model evaluates HONO sources, sinks in outflow of Dallas-Fort Worth
- Monte Carlo simulation is applied to scenarios with 3 recently identified sources
- Improved model outcomes result from inclusion of 2 of 3 recently identified sources
- A substantial unknown source is still required for agreement with observation
- Missing HONO source is moderately correlated with  $j_{NO2}$ , weakly correlated with  $NO_2$